

Due *in class* Thursday April 11th

Central Force Motion:

1. *K&K* Problem 10.2, *except* do it in MKS units with $m = 0.075$ kg, force = $80 r^3$ Newtons and $L = 2.0 \times 10^{-4}$ kg \cdot m²/s.
2. *K&K* Problem 10.4.
3. *K&K* Problem 10.6.
4. *K&K* Problem 10.9.

1) K&K 10.2

A particle of mass $m = 0.075 \text{ kg}$ moves under an attractive central force of magnitude $80r^3 \text{ Newtons}$. The angular momentum is equal to $2 \times 10^{-4} \text{ kg m}^2/\text{s}$.

$$a) U_{\text{eff}} = \frac{L^2}{2mr^2} + U(r)$$

$$-\frac{dU(r)}{dr} = F(r)r^{\hat{1}} = -\left(\frac{80 \text{ N}}{\text{m}^3}\right) \cdot r^3 = -Ar^3$$

$$\text{so } U(r) = \frac{20 \text{ N}}{\text{m}^3} r^4$$

$$\text{And } \frac{L^2}{2m} = \frac{(2 \times 10^{-4} \text{ kg m}^2/\text{s})^2}{2 \times 0.075 \text{ kg}} = 2.67 \times 10^{-7} \text{ Joule} \cdot \text{m}^2$$

$$\text{so } U_{\text{eff}} = \frac{2.67 \times 10^{-7} \text{ Joule} \cdot \text{m}^2}{r^2} + \frac{20 \text{ Joule}}{\text{m}^4} \cdot r^4$$

b) Circular motion occurs at the bottom of U_{eff} curve

$$U_{\text{eff}} = \frac{L^2}{2mr^2} + \frac{1}{4} Ar^4$$

$$\left. \frac{dU_{\text{eff}}}{dr} \right|_{r=r_{\text{circ}}} = 0 = -\frac{L^2}{mr_{\text{circ}}^3} + Ar_{\text{circ}}^3$$

$$r_{\text{circ}} = \left(\frac{L^2}{mA} \right)^{1/6} = \left(\frac{2 \cdot 2.67 \times 10^{-7} \text{ J} \cdot \text{m}^2}{4 \cdot 20 \text{ J/m}^4} \right)^{1/6}$$

$$r_{\text{circ}} = 0.04338 \text{ m}$$

And

$$U_{\text{eff}}(r_{\text{circ}}) = 0.02125 \text{ mJ}$$

< See Attached Figure >

↳ This is total energy for a circular orbit.



$$1c) \quad U_{\text{eff}}(r_0) = U_{\text{eff}}(2r_0)$$

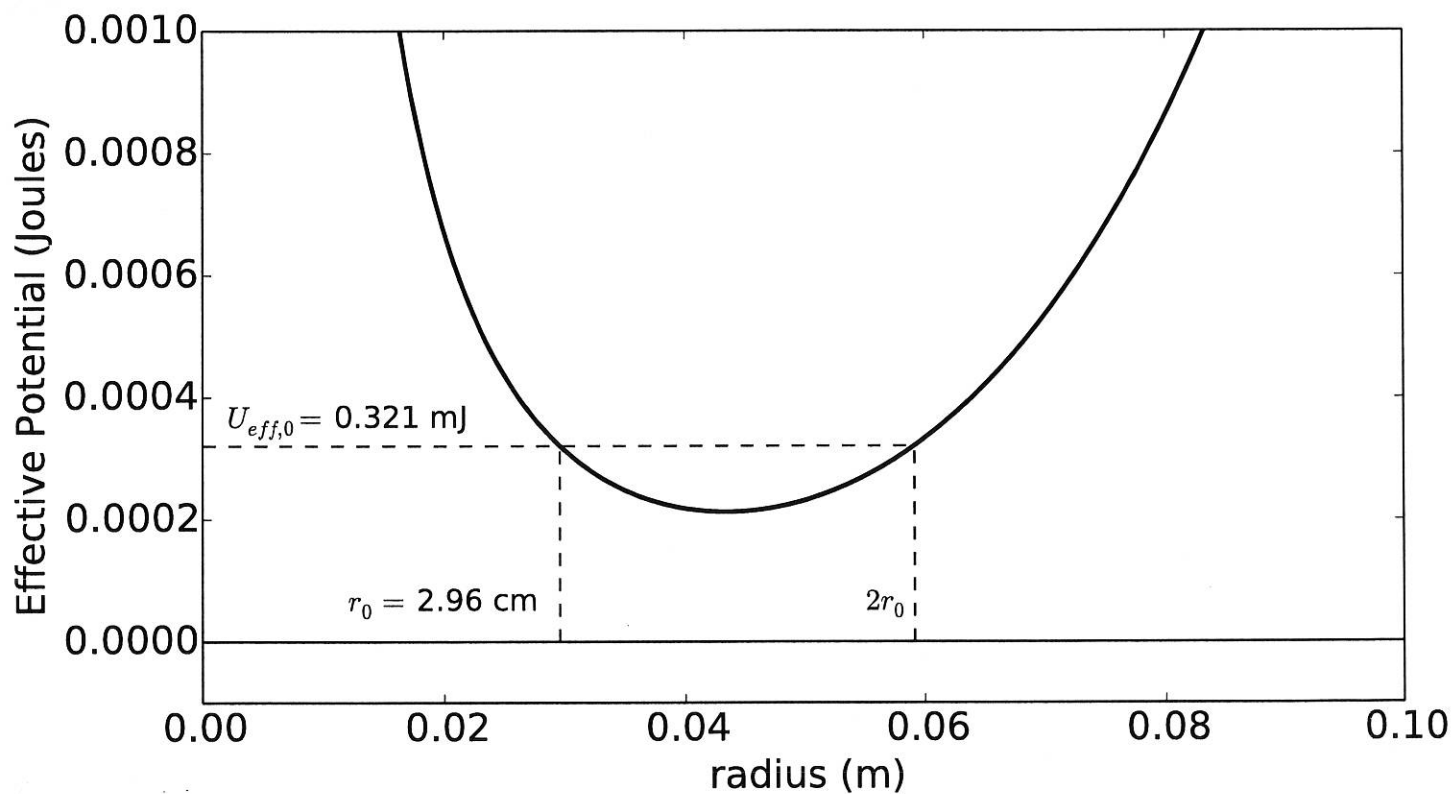
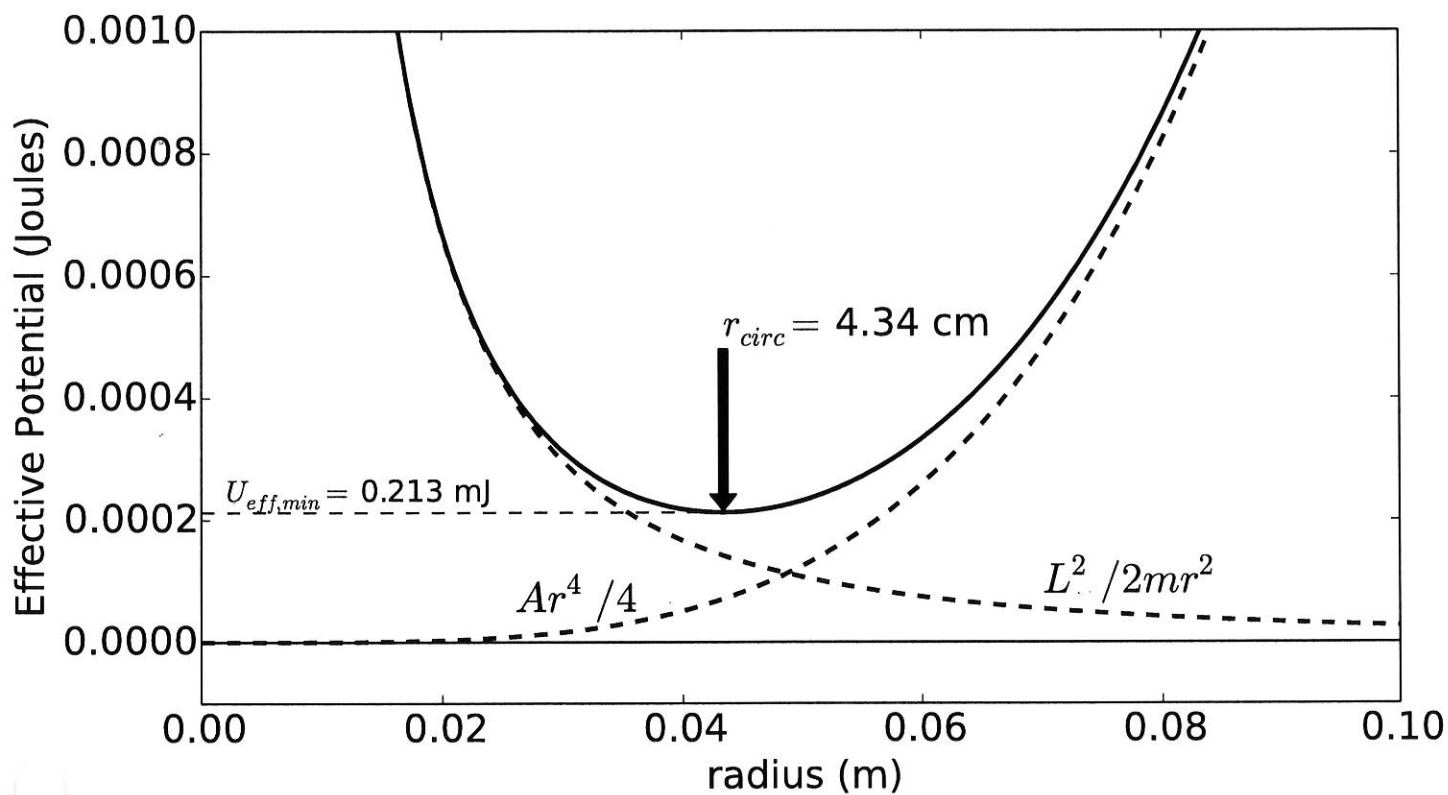
$$\frac{L^2}{2mr_0^2} + \frac{1}{4}Ar_0^4 = \frac{L^2}{2m(2r_0)^2} + \frac{1}{4}A(2r_0)^4 = \frac{L^2}{8mr_0^2} + 4Ar_0^4$$

$$\frac{L^2}{2m} + \frac{1}{4}Ar_0^6 = \frac{L^2}{8m} + 4Ar_0^6$$

$$\frac{3}{8} \frac{L^2}{m} = \frac{15}{4} Ar_0^6$$

$$r_0 = \left(\frac{1}{10} \frac{L^2}{m\hbar^2} \right)^{1/6} = \boxed{2.96 \text{ cm}}$$

< See Attached Figure >



2) K&K 10.4

$$U_{\text{eff}} = \frac{L^2}{2mr^2} - \frac{A}{r^n}$$

THE CONDITION FOR A CIRCULAR ORBIT TO EXIST IS THAT THERE BE A MINIMUM TO U_{eff} AT SOME FINITE RADIUS r_0 :

$$\left. \frac{dU_{\text{eff}}}{dr} \right|_{r=r_0} = -\frac{L^2}{mr_0^3} + \frac{nA}{r_0^{n+1}} = 0$$

$$\frac{nA}{r_0^{n+1}} = \frac{L^2}{mr_0^3} \quad (i)$$

$$\left. \frac{d^2U_{\text{eff}}}{dr^2} \right|_{r=r_0} = \frac{3L^2}{mr_0^4} - \frac{n(n+1)A}{r_0^{n+2}} > 0 \quad (ii)$$

SUBSTITUTING (i) INTO (ii)

$$\frac{3L^2}{mr_0^4} - \frac{(n+1)L^2}{mr_0^4} > 0$$

$$3 > n+1$$

$$\text{OR } \boxed{n < 2}$$

One problematic case would be $n=0$, which we can see from (i) would imply an infinite radius.

$$\hookrightarrow \boxed{n < 2, \text{ excluding } n=0 \text{ has circular orbit}}$$

3) K & K 10.6

$$\vec{F} = -K r^4 \hat{r} = -\frac{dU}{dr}, \quad U = \frac{1}{5} K r^5$$

$$U_{\text{eff}} = \frac{L^2}{2mr^2} + \frac{1}{5} K r^5$$

Circular motion will occur at minimum point on U_{eff} curve:

$$\left. \frac{dU_{\text{eff}}}{dr} \right|_{r=r_0} = -\frac{L^2}{mr_0^3} + Kr_0^4 = 0$$

$$r_0 = \left(\frac{L^2}{Km} \right)^{1/7}$$

$$\begin{aligned} E = U_{\text{eff}}(r_0) &= \frac{L^2}{2m} \left(\frac{Km}{L^2} \right)^{2/7} + \frac{1}{5} K \left(\frac{L^2}{Km} \right)^{5/7} \\ &= \boxed{\frac{7}{10} \left(\frac{L^{10} K^2}{m^5} \right)^{1/7}} \end{aligned}$$

Frequency of small oscillations: Perform Taylor Series Expansion of U_{eff} about $r=r_0 \dots$

$$U_{\text{eff}}(r) = U_{\text{eff}}(r_0) + \underbrace{\frac{dU_{\text{eff}}(r_0)}{dr}}_{=0} \cdot r + \frac{1}{2} \underbrace{\frac{d^2U_{\text{eff}}(r_0)}{dr^2}}_{\text{"effective spring constant" } k} \cdot r^2 + \dots$$

$$k = \left. \frac{d^2U_{\text{eff}}}{dr^2} \right|_{r=r_0} = \left(\frac{3L^2}{mr_0^4} + 4Kr_0^5 \right) = 7K \left(\frac{L^2}{Km} \right)^{3/7} \quad (\text{ANIMATE SKIPPED})$$

$$\boxed{\omega = \sqrt{\frac{k}{m}} = \sqrt{\frac{7K \left(\frac{L^2}{Km} \right)^{3/7}}{m}}}$$

4) K & K 10.9 Halley's Comet

- a) Determine perhelion & aphelion distances
(r_{\min} & r_{\max} in figure pg 391).

$$r_{\text{per}} = \frac{r_0}{1 + e}$$

$$r_{\text{ap}} = \frac{r_0}{1 - e}$$

Eq 10.28 $A = r_{\text{per}} + r_{\text{ap}} = r_0 \left(\frac{1}{1+e} + \frac{1}{1-e} \right) = \frac{2r_0}{1-e^2}$

Eq 10.31 $T^2 = \frac{4\pi^2 m}{2G} A^3$

$$\therefore A = \left(\frac{2T^2 G M_{\text{sun}}}{\pi^2} \right)^{1/3} = 5.37 \times 10^{12} \text{ meters}$$

and

$$r_0 = \frac{1}{2} A (1 - e^2) = 1.74 \times 10^{11} \text{ m}$$

$$\therefore \begin{aligned} r_{\text{per}} &= \frac{r_0}{1+e} = 8.86 \times 10^{10} \text{ m} \\ r_{\text{ap}} &= \frac{r_0}{1-e} = 5.27 \times 10^{12} \text{ m} \end{aligned}$$

- b) What is speed at perhelion?

Eq 10.19 $r_0 = \frac{L^2}{m G M_{\text{sun}} m} = \frac{m^2 v_{\text{per}}^2 r_{\text{per}}^2}{m G M_{\text{sun}} m}$

$$v_{\text{per}} = \left(\frac{G M_{\text{sun}} r_0}{r_{\text{per}}^2} \right)^{1/2} = 5.42 \times 10^4 \text{ m/s}$$